

FROM HERE AND THERE -- Cont'd. from p. 3

We would like to print a list in STARDUST of the NCA members that belong to other astronomical groups such as the AAVSA, ALPO, the AMS and others. If you belong to one of these groups, please give the information to the secretary, Mrs. Nora Keel.

NEXT MONTH

Our speaker for March will be Dr. David Musto, who will describe the beginnings of the U.S. Naval Observatory.

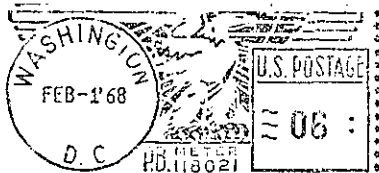
Astronomy compels the soul to look upwards
And leads us from this world to another.

- Plato

NO MD-DC JUNIOR MEETING in FEBRUARY.

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Library,
Naval Observatory
Washington 25, D.C.



★ S T A R D U S T



February 1968

Vol. XXV No. 6

GRAVITY WAVES

February's speaker is Dr. Joseph Weber, from the Department of Physics and Astronomy, University of Maryland. Dr. Weber's area of investigation recently has been gravity waves; he is in charge of experiments which promise possible detection of this elusive form of radiation.

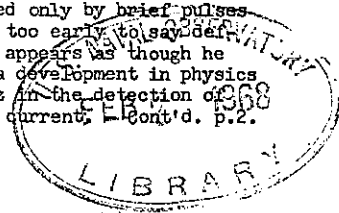
Relativity predicts that an oscillating mass will radiate energy in the form of gravity waves, much as an alternating current will produce electromagnetic radiation. However, this radiation is on the order of only 10^{-37} watts for a terrestrial source*, making detection virtually impossible. Even a spectroscopic binary star system does not radiate sufficient energy to be detected by means we now have at our disposal. Nature may provide a much more intense source of gravity waves; this might possibly occur when a star undergoes gravitational collapse. Conservation of angular momentum in the system would dictate faster and faster rotation as the radius shrinks, until the velocity of light is approached. Gravitational radiation energy of perhaps 10^{50} watts would be attained. This could very well be detected with equipment now available to the physicist.

Dr. Weber's apparatus, which he will describe in his talk, has detected "events" which can apparently be explained only by brief pulses of gravity radiation, although it is still much too early to say definitely whether this is the case. Certainly it appears as though he and his staff are onto something big, possibly a development in physics fully as significant as the experiments of Hertz in the detection of electromagnetic waves emitted by an alternating current. Cont'd. p.2.

CALENDAR

- FEBRUARY 3 6:00 p.m. Dinner with the speaker, Bassin's (14th & Pa.) Call Jerry Hudson, 948-2809 for reservations.
- 7:30 p.m. meeting of the officers and board of trustees. (Because of bad weather, we were unable to meet a quorum last month) Dept. of Commerce Auditorium.
- 8:15 p.m. GRAVITY WAVES, Dr. Joseph Weber, Univ. of Maryland. Dept. of Commerce Auditorium.
- 10 8:15 p.m. Discussion Group. (upstairs) Commerce Bldg., Rm.2062. Topic will be an amateur's observatory--that of our treasurer, Mr. Robert N. Bolster. Slides will be shown which were taken during various phases of construction, and mention will be made of certain Fairfax County zoning regulations pertinent to observatory building.
- 6,13,20,27 TELESCOPE MAKING CLASS at the Board of Education's Material Center in Bladensburg with Ted Noble.
- 2,9,16,23 TELESCOPE MAKING CLASS at the Chevy Chase Community Center at 7:30 P.M. with Hoy Walls.
- 3 GENERAL MEETING OF JUNIOR DIVISION at 7:30 P.M. in the Dept. of Commerce auditorium. All juniors are urged to attend. January Junior Meeting was cancelled due to snow.
- JUNIOR DIVISION ASTRONOMY CLASS Sunday afternoon at 3 P.M. in Planetarium of Montgomery Junior College. Call 933-0823 (Mark Goldberg) for the date.

* This was calculated by Einstein for a rotating rod 1-meter long, at an angular velocity as fast as the material would allow.



QB
1
S79

GRAVITY WAVES - Cont'd. from p. 1

Dr. Weber received his Ph.D. in physics from the Catholic University of America in 1951. He has been the recipient of two Guggenheim fellowships under which he has done work in quantum electronics and general relativity. Before coming to the University of Maryland, he was with the Institute for Advanced Studies.

Those who may want to do some more reading on the subject will find an excellent, non-technical article in the October 1967 issue of Scientific Research (pp. 41-43).

MIDDLE EAST REGIONAL CONVENTION JUNE 14, 15, and 16, Pittsburgh, Pa.

The convention will be held at the Quality Courts Motel East (near the Westinghouse Bridge). It is on Route 30 off the Pa. Turnpike at the Irwin interchange. The motel is giving special rates. Friday evening will be a bus trip to the Allegheny Observatory. The Banquet will be Saturday evening with Dr. Wagman of the Allegheny Observatory being the speaker. Sunday will be a trip to large private Observatory in the Pittsburgh area. Plan now to attend this convention. In case you will consider presenting a paper, please write to Mr. George G. Lingbloom, 1606 Burchfield Road, Allison Park, Pa. 15101.

NEW MEMBERS

Membership Applications Received at the January Meeting

Joint Membership:

Darrel J. Freund and Darrel Jr.
4703 Teak Ct.
Camp Springs, Md. 20031

NOTE FROM EDITOR

All of us associated with the publication of Stardust regret that it is often received late and sincerely hope that this has not caused you to miss a lecture that you particularly wanted to hear. The publication of Stardust requires the efforts of many people to collect the information necessary to write even just the calendar! It should be remembered that these people, including the editor, and the photographer are all voluntary and can not be expected to take time away from their regular jobs to work on Stardust. If you have never worked on a project like this you may not believe the number of things small and large that can go wrong--a last minute change in speaker, the speaker forgets to send his resume, someone is sick, someone is called out of town. When an emergency arises, there is no one prepared to take over and fill the gap. Several times in the last two years requests have been made in Stardust and at the regular meetings for help. No one has volunteered! Stardust belongs to the membership and your help is needed to make it the type of publication you want.

FROM HERE AND THERE

The Astronomical Society of Harrisburg will hold their annual dinner February 15, Gallileo's birthday. Dr. Peter Van De Kamp will be the speaker.

How about the NCA having an annual dinner? Guess it would be our 31st year. Seems like it has been 10 or 12 years since we have had a banquet. Someone get the ball rolling.

We also note in the "STARDUST" of the Harrisburg group that their Junior group is selling the Messier OBSERVE Manual for \$1.50. Copies are still available from Bob Wright for \$1.00.

- Cont'd.p. 3

LUNAR OCCULTATIONS COMPUTED FOR WASHINGTON (LAT 38.920 LONG 77.065) FOR 1968 (DISTANCE FROM STANDARD STATION = -0 MILES) STANDARD STATION AT LAT 38.920 LONG 77.065

Table with columns: DATE, TIME(UT), ACC, V O Z.C., USNO REF NO, S.A.O. NUMBER, MAG, SP, PH, PERCENT SUNLIT, ELG, PA, DEG ANGL, CUSP VA, DEC, MOON ALT AZI, SUN ALT, LIBRATION DEG LONG, MIN LAT MAG.

OCCULTATION PREDICTIONS

These predictions are provided by David W. Dunham (Yale Observatory) and Thomas C. Van Flandern (U.S. Naval Observatory). The following explanations apply to column headings:

TIME (UT): Note that events are generally, but not always, in strict chronological order.

ACC SEC: Approximate accuracy of predicted time, in seconds.

V: Value code, indicating expected usefulness of the observation.

O: Observability code, indicating expected difficulty/ease. 9= easiest.

Z.C.: Zodiacal Catalog number. A=double star (Aitken's catalog). B or C= other double star. D= double and variable star.

S.A.O. NUMBER: Star's number in Smithsonian Astrophysical Obs. catalog.

PH: Phenomenon. G= graze. M=miss. It is possible for extremely shallow solid occultations to be listed as grazes, and vice-versa. ELG: Elongation, i.e., distance of Moon from Sun, in degrees. PA DEG: Position angle of the event, in degrees, meas. eastward from the north point of Moon's disc (not from the north cusp). CUSP ANGL: Distance of event meas. onto dark limb from N or S cusp. VA DEG: Vertex angle, meas. counterclockwise from uppermost (toward the zenith) point on Moon's disc (for use with alt-azimuth instruments). SUN ALT: Given only when critical. AA DEG: Axis angle, meas. eastward from Moon's north pole. MIN MAG: Magnitude at min. brightness of var. star (for which MAG=max.) GRAZING OCCULTATION NEARBY--<cc.: This extra line is printed out if the star is within 0.02 radius (approx. 20 miles) of the mean limb at central occultation or closest approach. (This line refers to prev. line).

LUNAR OCCULTATIONS COMPUTED FOR WASHINGTON (LAT 38-920 LONG 77-065) FOR 1968 (DISTANCE FROM STANDARD STATION = -0 MILES) STANDARD STATION AT LAT 38.920 LONG 77.065

Table with columns: DATE, TIME(UT), ACC, V, O, Z.C., S.A.O. NO, MAG, SP, PH, PERCENT, ELG, CUSP ANGL, MOON SUN AZ, LIBRATION, MIN. Contains multiple rows of occultation data for various stars.

JANUARY LECTURE - ROCKETS, ASTRONOMY, AND AERONOMY

X-ray stars are the hottest topic in astronomy today according to our January speaker--Mr. Charles Johnson-- head of the Aeronomy Section of the Naval Research Laboratory. The most famous X-ray source is the Crab Nebula (Taurus XR-1) which an Aerobee rocket flight proved several years ago is not a point source of X-rays. This flight occurred during an occultation of the nebula by the moon, and the X-ray emission did not cut off sharply but decreased gradually. The strongest X-ray source is Scorpius XR-1. Most X-ray sources exist along the galactic equator with the greatest concentration toward Sagittarius. Few are extra-galactic; a notable example of one of these is quasar 3C-273. One variable X-ray star has been discovered; i.e., Cygnus XR-1, whose strength decreased significantly between June 1964 and the following April.

Aeronomy is a new field of science barely a decade old which deals with that part of the upper atmospheres of planets where ionization and disassociation are important. Here the density of air is millions of times less than at the surface.

The German V2 rockets were the first to observe and record the ultra-violet spectrum of the sun. Only by rocket or satellite can the sun's Lyman alpha line of hydrogen (1216 A.) be seen. After the V2's had been used up, N. R. L. Aerobee rockets continued upper atmosphere research by carrying a wide variety of instruments into near space up to 250 kilometers. Various bands of ultraviolet can be selected and measured by detectors having different combinations of windows and gases; for example, nitric oxide exposed to uv through a lithium fluoride window. The window sets a lower limit on the detected radiation wavelength and the gas, an upper limit. A Schmidt camera with a calcium fluoride corrector plate and image intensifier electronics has been used on board Aerobees for studying hot O and B type stars in ultraviolet light down to 1250 A. These optics are blind to Iyman alpha radiation which is strong in air glow and would fog uv star fields.

The Bennett R F mass spectrometer has been successfully used on Aerobees for measuring ion concentrations in our upper atmosphere. This instrument has revealed many interesting facts about the ionosphere such as the presence of a great deal of the nitrous oxide ions in our upper atmosphere. (Perhaps there is a Martian Dr. Kiess who thinks that life is impossible on the earth because of the presence of this poisonous compound of nitrogen in our atmosphere!) The Marina V fly-by of Venus studied the ionosphere of this planet. Apparently the ionosphere of Venus is high at night but low during daytime. Mr. Johnson explained that since Venus has no appreciable magnetic field, the solar wind is strong enough to blow the ionosphere off the planet or to the night side.

- Leith Holloway

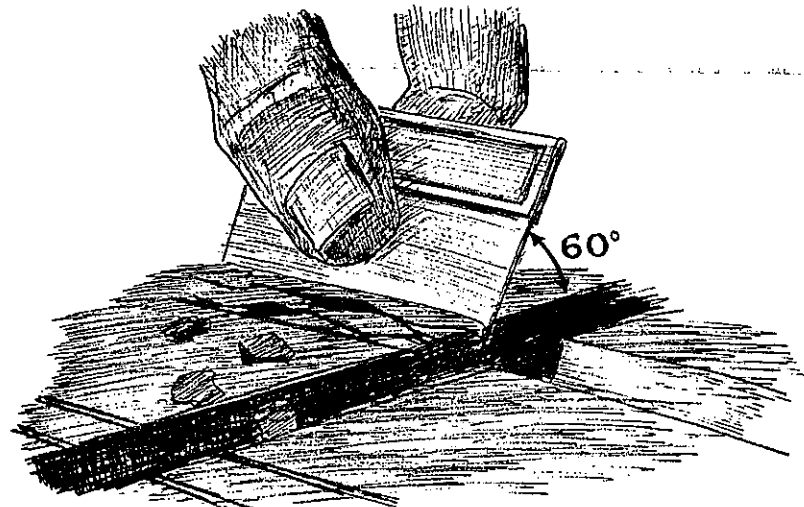
FROM HERE AND THERE...Cont'd. from p.2

From "THE GUIDE STAR" of the AAA of Pittsburg: How are we going to teach logic in a world where everybody talks about the Sun setting, when it is really the horizon rising?

"STAR LITE" is the name of the newsletter sent out by the Peoria Astronomical Society of Peoria, Ill.

"STAR Newsletter" is the title of the publication of the Society of Telescopy, Astronomy and Radio (STAR) of Eatontown, N.J., a society belonging to the Middle East Region of the Astronomical League.-Concluded on p. 8

The January-February issue of Review of Popular Astronomy carries an article by Jerry Hudson. It is the same article that appeared in a recent STAR DUST on A Simple Clock Drive Design.



Before one has gone once around the barrel, the HCF has worn into contact with the optical surface.

HCF is useful too when making gross zonal corrections, such as removing a severe case of turned-down edge, or correcting a hyperbola. Final smoothing and blending can only be done on a pitch lap, however.

This suggests one might wish to use another piece of glass (preferably one with approximately the right curvature) for the pitch lap, keeping the original tool free for HCF working. One can proceed a bit more confidently on his pitch lap, knowing he has an intermediate alternative to the one dreaded by all mirror-makers returning to find grinding.

OCCLUSION PREDICTIONS

The use of the table is self-explanatory, with possibly the following explanations needed:

Z.C.= Zodiacal Catalogue number. If followed by "M", take the mean time of the appearance (disappearance) of the double star. For any other letter, use the brighter component of the pair.

Some Z.C. numbers have been assigned to bright Solar System objects:

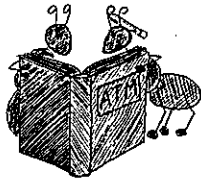
- Mercury..... 4001
- Venus..... 4002
- Mars..... 4004
- Jupiter..... 4005
- Saturn..... 4006
- Uranus..... 4007
- Neptune..... 4008
- Ceres..... 5001
- Pallas..... 5002
- Juno..... 5003
- Vesta..... 5004

PHEN= Phenomenon D= Disappearance R=Reappearance =Near Grasser
 P.A. DEG=Position Angle (degrees), measured eastward from the northernmost point on the moon's disc. For example, 90° would be the center of the dark limb during waxing; 270° during waning.

P.A.C.B.L.=Position Angle of the Center of the Bright Limb.(same convention)
 Altitudes of the sun and moon are only given when these are important.

LUNAR OCCULTATIONS COMPUTED FOR WASHINGTON
 (LAT. 38.920 LONG. 77.065) FOR 1968 (DISTANCE FROM STANDARD STATION = 0 MILES)
 STANDARD STATION AT LAT 38.920 LONG 77.065

DATE	TIME(L) M H S	ACC	V	D	Z.C.	USNO REF NO	S.A.G. NUMBER	MAG	SP	PH	PERCENT	ELG PA	CUSP VA	MOON DEC ANGL	SUN ALT	LIBRATION HI		
JAN 2	22 42 50	6	7	9	3143A	223031	190356	7.5	K0	D	10	WAX	36 22	33N 347	18 223	-9 39	5.6 6.6	
JAN 3	22 36 53	10	5	8		223006	165002	8.9	K2	D	17	WAX	AR 113	80S 16	25 214	-11 133	5.8 5.8	
JAN 3	23 47 40	5	5	6		223925	165016	8.6	G5	D	17	WAX	AR 17	54N 359	20 229	57	5.6 5.7	
JAN 3	23 48 50	5	5	6		223935	165017	9.0	K0	D	17	WAX	AR 56	73N 18	20 230	76	5.6 5.7	
JAN 4	0 19 14	5	6	5		223947	165026	8.2	G5	D	17	WAX	AR 62	79N 21	15 235	83	5.6 5.6	
JAN 4	1 11 46	39	6	9	328A	223961	165032	7.1	F5	D	17	WAX	AR 340	135S 95	7 244	160	5.5 5.5	
JAN 4	1 19 56	38	6	9	376A	223961	165032	7.1	F5	R	17	WAX	AR 155	75 108	5 246	175	5.5 5.5	
JAN 4	22 36 26	11	4	9		224584	166509	7.3	K0	D	25	WAX	AR 368	9N 328	38 205	-7	9 5.6 4.8	
JAN 4	23 1 13	5	5	7		224609	166517	8.8	G0	D	25	WAX	AR 61	62N 16	36 211	-12	62 5.5 4.7	
JAN 4	22 51 18	5	5	8		224610	166518	8.9	F0	D	25	WAX	AR 80	81N 38	37 203	-10	82 5.5 4.8	
JAN 5	0 22 47	5	4	7		224631	166536	8.6	G0	D	25	WAX	AR 60	59N 35	25 231	54	5.3 4.6	
JAN 5	1 21 16	8	6	6		224649	166547	9.1	K5	D	26	WAX	AR 61	104S 55	52 17 243	125	5.2 4.5	
JAN 5	1 49 3	6	6	7		224660	166556	8.6	F8	D	26	WAX	AR 61	89 70S	42 12 248	111	5.1 4.4	
JAN 6	0 34 7	5	6	7		225232	166572	8.8	F8	D	32	WAX	AR 72	62N 27	35 228	85	4.6 3.4	
JAN 6	2 22 2	5	7	7		225273	166992	8.8	F8	D	35	WAX	AR 73	55 78N	6 18 251	77	4.3 3.2	
JAN 6	3 48 38	8	7	7		225307	167015	7.3	F0	D	36	WAX	AR 73	110S 58	52 245	132	4.2 3.0	
JAN 6	22 29 22	12	5	6		200462	109303	9.0	F0	D	44	WAX	AR 83	112	45S 116	52 176	-6	134 4.1 2.6
JAN 6	23 20 52	9	5	6		200470	109310	9.0	MA	D	44	WAX	AR 83	106S 51S	93 52 196	127	3.9 2.2	
JAN 7	0 33 45	5	6	8		200483	109319	9.1	F8	D	44	WAX	AR 83	66N 12	45 222	65	3.6 2.0	
JAN 7	1 3 43	6	6	8		200493	109325	9.0	ER	D	44	WAX	AR 83	17	60N 33	41 230	38	3.2 2.0
JAN 7	0 55 39	5	6	6		200502	109335	8.9	F5	D	44	WAX	AR 83	57	80N 21	42 228	78	3.6 2.0
JAN 7	3 33 3	5	7			200565	109385	8.8	K2	D	45	WAX	AR 84	83	75S 32	16 261	104	3.2 1.7
JAN 7	3 34 33	7	6	6		200566	109386	9.0	D	D	45	WAX	AR 84	105S 52S	55 16 261	127	3.1 1.7	
JAN 7	4 25 20	7	7			200587	109399	8.8	F5	D	45	WAX	AR 85	19N 42N	32 28	6 269	41	3.1 1.7
JAN 7	22 28 10	5	9			201139	109402	7.8	G5	D	53	WAX	AR 94	70	88S 88	56 157	-5	91 2.9 0.9
JAN 7	22 58 38	6	7			201147	109808	8.2	G5	D	53	WAX	AR 94	74S 92	38 176	-11	104 2.8 0.8	
JAN 8	1 10 23	5	7			201183	109827	8.6	G5	D	54	WAX	AR 94	80S 86	51 225	84	2.3 0.6	
JAN 8	2 40 16	8	5	9		201217	109853	8.2	A3	D	56	WAX	AR 95	109S 49S	62 38 248	129	2.1 0.4	
JAN 8	4 45 20	7	6	7		201272	109894	8.6	K2	D	55	WAX	AR 96	58	13 270	131	1.8 0.3	
JAN 8	5 18 2	6	8	6		201286	109906	8.9	K0	D	55	WAX	AR 96	96S 44S	42 7 275	143	1.7 0.3	
JAN 8	22 12 12	5	6	2		201794	92761	7.5	K0	D	63	WAX	AR 104	66N 103S	55 131	-2	84 1.7 -0.4	
JAN 9	1 13 43	14	6	7	365	201836	92788	8.7	K0	D	63	WAX	AR 105	125 36S	90 216	143	1.1 -0.8	
JAN 9	1 14 35	5	6	9		201838	92789	8.1	F5	D	63	WAX	AR 105	64	94N 35	60 217	82	1.1 -0.8
JAN 9	1 51 35	5	6	9		201848	92795	7.9	K0	D	64	WAX	AR 106	64	85N 25	55 231	82	0.9 -0.9
JAN 9	4 35 23	6	7	7		201889	92814	8.7		D	64	WAX	AR 106	61	94N 34	26 267	59	0.4 -1.0
JAN 9	5 33 8	5	7	7		201891	92816	8.8		D	64	WAX	AR 106	45	69N 35S	27 267	67	0.4 -1.0
JAN 9	6 11 4	12	9	9	328	201925	92837	7.9	K0	D	65	WAX	AR 107	136	25S 84	8 201	154	0.3 -1.1
JAN 9	6 16 33	5	7			201929	92840	8.7	K0	D	65	WAX	AR 107	57	76N 6	7 282	76	0.3 -1.1
JAN 9	22 25 93	5	7	5		202327	93112	9.2		D	72	WAX	AR 115	63	60N 90	54 117	-5	94 0.4 -1.6
JAN 9	22 31 19	5	3	9		202331	93115	7.4	M6	D	72	WAX	AR 115	77	86S 124	55 118	-6	93 0.4 -1.8
JAN 10	3 47 2	10	8	8		202388	93167	8.8		D	73	WAX	AR 117	6	22N 314	47 255	22	0.7 -2.3
JAN 10	4 42 25	15	9	9	432	202419	93189	9	M8	D	73	WAX	AR 117	141	23S 86	35 265	156	-0.9 -2.4
JAN 10	6 40 25	5	6	6		202466	93225	9.2		D	74	WAX	AR 118	95	69S 42	14 283	110	-1.1 -2.4
JAN 10	7 33 42	5	5	5		202473	93230	9.0		D	74	WAX	AR 118	166	-2S 116	4 291	181	-1.2 -2.4
GRADING OCCULTATION NEARBY - APPROXIMATE S. LIMIT LAT. = 32.04 LONG. = 77.06																		
JAN 10	22 48 13	5	7	4		202921	76022	8.8	G5	D	80	WAX	AR 127	52	64N 126	52 106	-8	64 -0.9 -3.1
JAN 10	22 12 12	8	7	3		202944	76027	8.1	G0	D	80	WAX	AR 127	117	87S 155	55 121	87	-2.5 -3.2
JAN 11	0 48 35	12	7	6		202960	76050	7.8	A0	D	80	WAX	AR 127	126	42S 150	71 152	138	-1.3 -3.3
JAN 11	2 20 23	11	7	3		202977	76067	8.8	A2	D	81	WAX	AR 128	3	14N 332	70 188	15	-1.6 -3.5
JAN 11	1 50 10	5	7	4		202979	76069	8.7	K2	D	80	WAX	AR 127	42	53N 27	73 197	54	-1.5 -3.4
JAN 11	4 39 29	22	8	4		203015	76097	8.6	F5	D	81	WAX	AR 128	358	9N 300	47 262	10	-2.1 -3.6
JAN 11	5 26 23	5	9	5	940	203070	76154	8.0	A0	D	81	WAX	AR 128	56	67N 358	38 270	68	-2.2 -3.6
JAN 11	6 58 58	5	9	6		203111	76156	8.6	A0	G	82	WAX	AR 129	351	2N 29S	21 283	2	-2.4 -3.6
GRADING OCCULTATION NEARBY - APPROXIMATE N. LIMIT LAT. = 38.93 + 0.341M LONG. = 77.06																		
JAN 11	7 10 46	5	9	6		203146	76224	8.4	A9	D	82	WAX	AR 129	64	75N 9	19 284	76	-2.4 -3.6
JAN 11	7 18 4	6	9	4		203154	76232	8.3	A0	D	82	WAX	AR 129	111	85S 57	17 285	123	-2.5 -3.6
JAN 11	7 40 50	10	6	8	563	203163	76242	8	BB	D	82	WAX	AR 129	141	29S 86	13 288	152	-2.5 -3.6
JAN 11	7 46 39	7	9	5		203175	76254	8.0	F5	D	82	WAX	AR 129	141	23S 86	35 265	156	-0.9 -2.4
JAN 11	7 49 52	5	9	3		203178	76257	8.7	F5	D	82	WAX	AR 129	105	65S 51	12 290	115	-2.2 -3.6
JAN 11	8 40 30	6	7	4		203206	76283	7.8	G0	D	82	WAX	AR 130	119	51S 7	3 297	131	-2.5 -3.4
JAN 11	22 58 1	6	7	4		203462	76612	8.2	A0	D	87	WAX	AR 138	37	43N 98	47 26	-10	45 -1.9 -4.2
JAN 12	2 27 6	19	7	1		203769	76637	8.7	G5	D	87	WAX	AR 138	143	67N 358	38 270	68	-2.2 -3.6
JAN 12	1 59 50	7	8	2	483	203715	76646	8.2	A2	D	88	WAX	AR 139	143	52S 129	74 165	124	-2.5 -3.6
JAN 12	5 31 34	6	4	6	698	203768	76676											



SHOP TALK

Page 4

POLARIS TELESCOPES

Many portable telescopes lack a ready means for lining up the polar axis with the earth's axis, thereby assuring the observer of hours of tracking by the clock drive, and making possible the taking of long exposures. Either the problem is ignored, whence observing becomes less pleasant and photographs unsatisfactory; or the observer spends half an hour tediously lining up the mounting, using his finder for sighting.

A glance of Fig. c shows why even careful guiding is not able to eliminate trailing of star images if the axis is misaligned. The film and camera remain in the same alignment north-south with respect to the false pole, F, while the direction on the film toward true north, P, changes during the exposure. Of course, the guide star remains fixed on the film, while all other parts of the field rotate about that point. The angle through which the film rotates is $\theta_1 - \theta_2$. Under the worst possible conditions, for a pole misalignment ϵ , a δ -declination of the guide star D ($90^\circ - \text{decl.}$), and an exposure of length T.

$$(\theta_1 - \theta_2)_{\text{max}} \approx \frac{\epsilon \omega T}{D}$$

ω is the angular rate of the earth's rotation, $15^\circ/\text{hr}$. Thus, if our polar axis is misaligned 2° , our guide star 30° from the pole, and we make an hour's exposure, we could rotate the field as much as 1° , which would easily be seen on any photograph. Cameras of short focal length are not immune to this effect--all that matter are film size and off-set of the guide telescope.

The polaris telescope shown in Fig. a is patterned after one built by Mr. Robert N. Bolster, and could be adapted to any standard mounting. A means for adjusting the alignment of the telescope is needed, such as the ring mounts shown. The reticle is a standard cross-line ruled on glass (obtainable from Edmund Scientific Co.). A ring of the correct angular radius ($54'$ for 1968.0, $+0'.3$ per ann.) is scribed onto stiff, transparent plastic with a pair of dividers. One should use very light pressure, and practice on several scraps before doing the final one. The plastic is then centered on the cross-line reticle, and the two mounted inside an eyepiece. The threaded retaining ring that was intended for the field lens can be used; machine washers can be inserted to obtain the right focus.

In use, the mounting is lined up approximately, and the main telescope pointed to any of the following bright stars:

- I. (0^h) α Andromedae (N.E. corner of the Great Square)
- II. (6^h) α Orionis (Betelgeuse)
- III. (12^h) γ Ursae Majoris (in the Big Dipper)
- IV. (18^h) γ Draconis (in the Dragon's head)

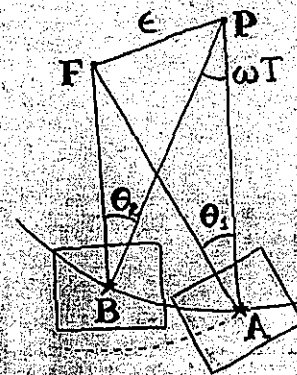
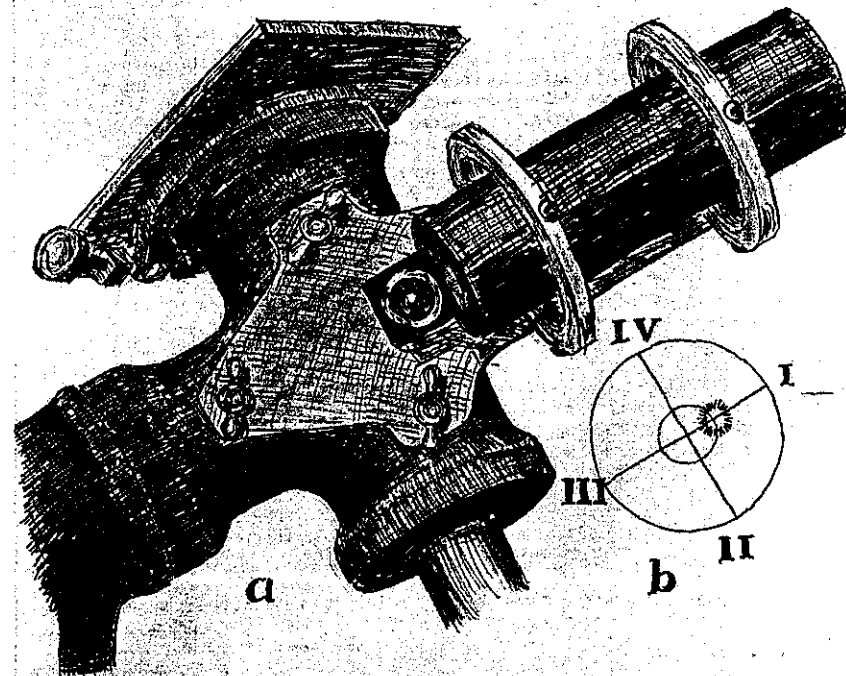
Final adjustment is made by lining up the Polaris telescope with Polaris on the appropriate hair and the scribed circle, as shown in Fig. b. The view is as one would look directly into the eyepiece in Fig. a, and it assumes a simple star diagonal was used in the Polaris telescope.

Adjusting the rig consists of bringing the intersection of the cross hairs into coincidence with the rotation of the mounting and turning the reticle until hair I coincides with an angle 30° (2^h) east of the meridian. The main telescope should be clamped to point along the meridian.

A little time in the shop making a Polaris telescope will be repaid during the first few nights of observing. Lining up should not take over five minutes, and accuracy can be held consistently to within $1/10$ degree.

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POLARIS TELESCOPE



Polaris
telescope

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